



Stockholm
University

Chromospheric Observation of OI (7772 Å)

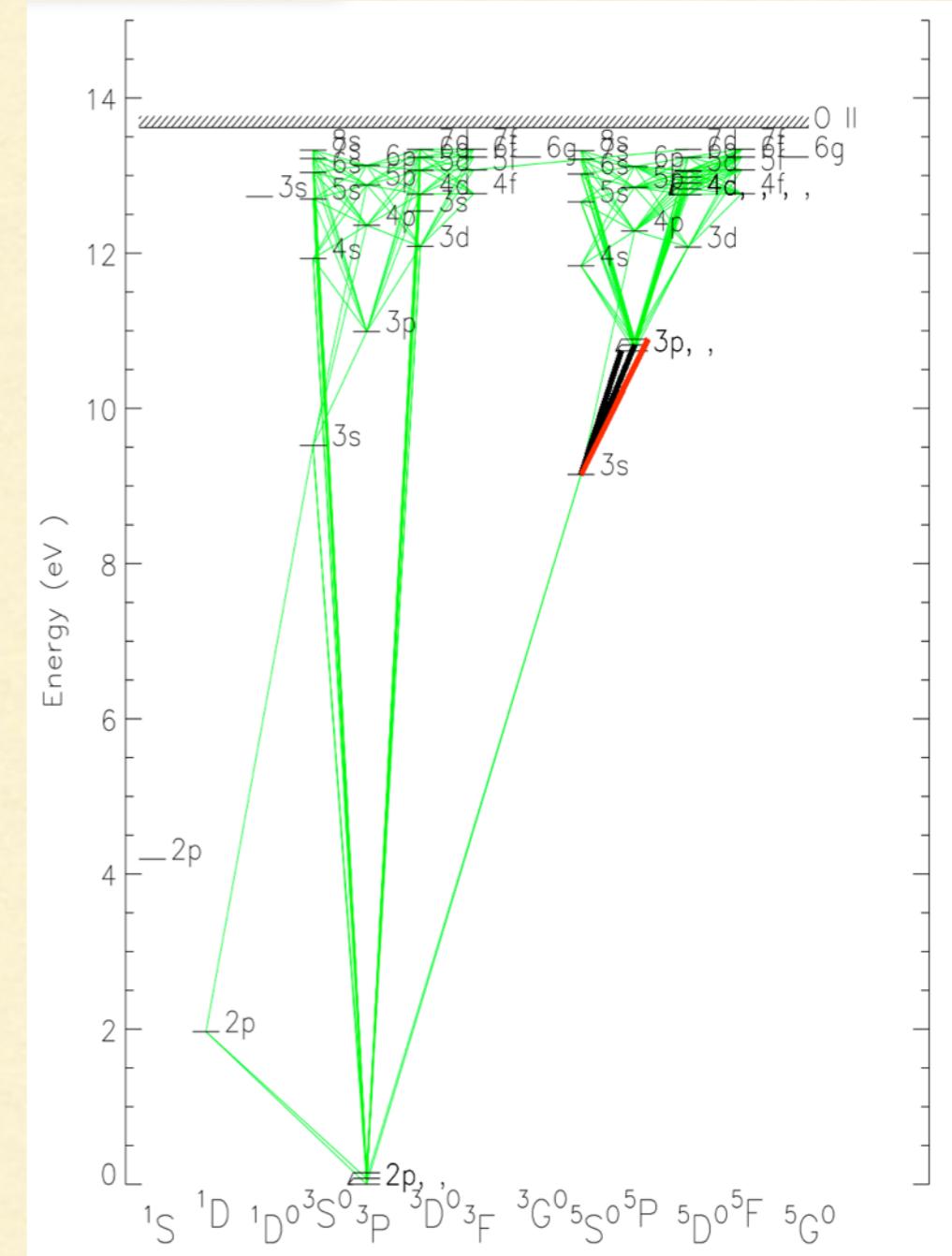
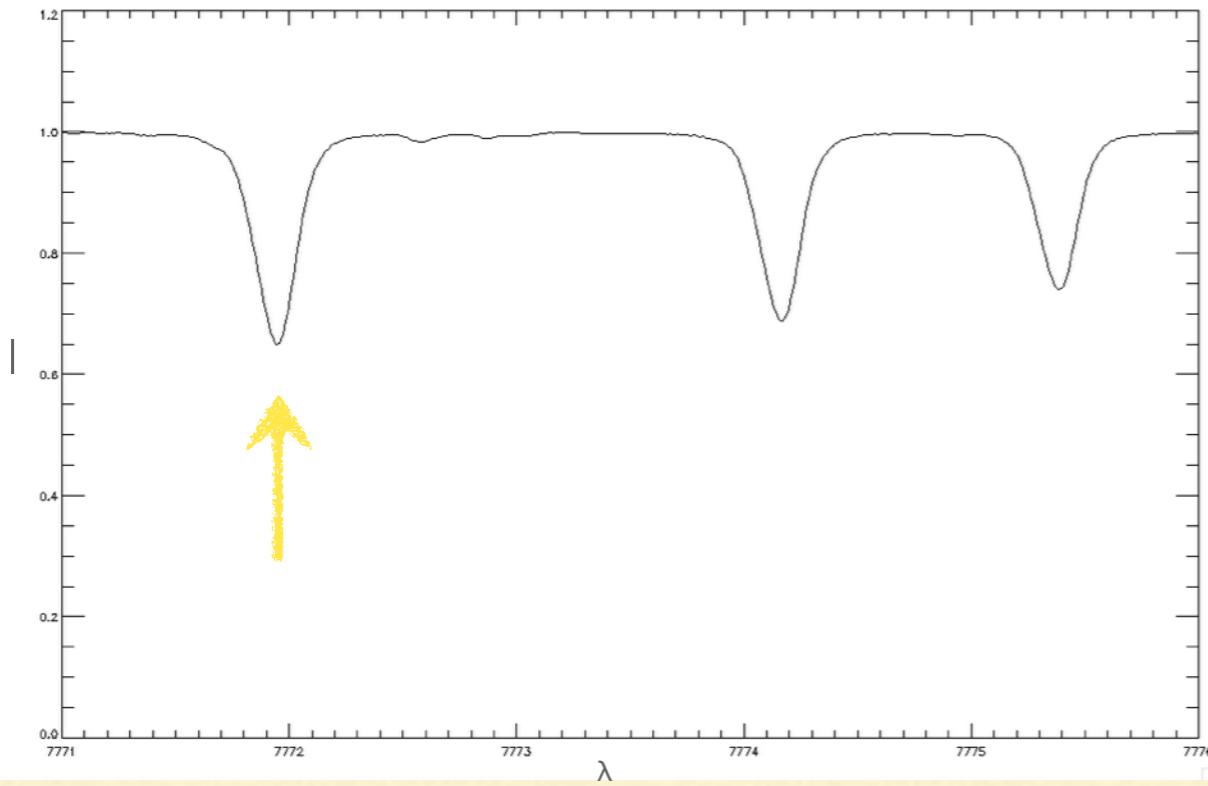
Hiva Pazira

Jaime de la Cruz Rodriguez, Dan Kiselman

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Why OI at 7772 Å

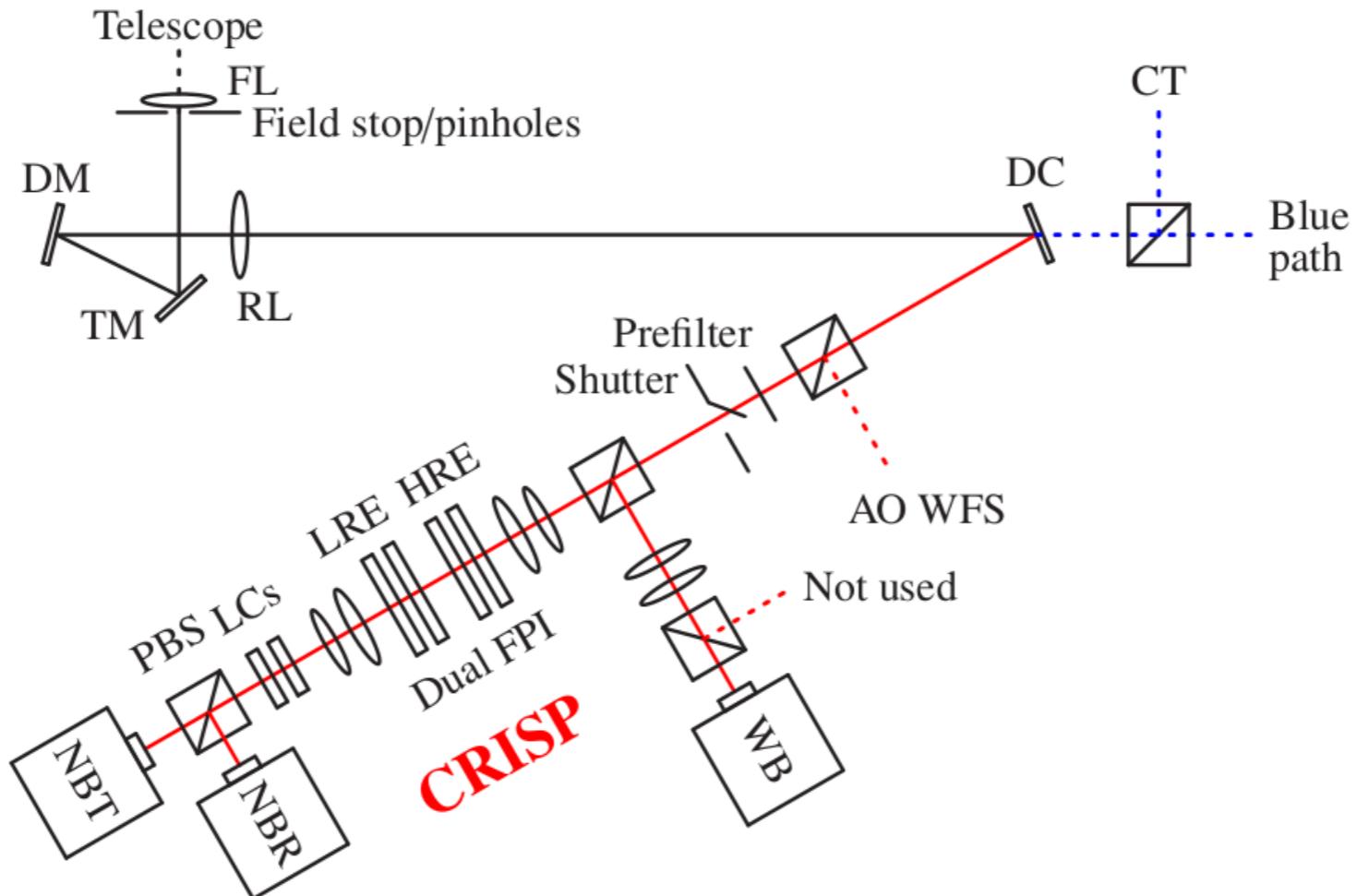
- A chromospheric line
- No blending
- Only observable in OI line in metal-poor stars
- No clear line formation process



Fabbian et al. (2009)

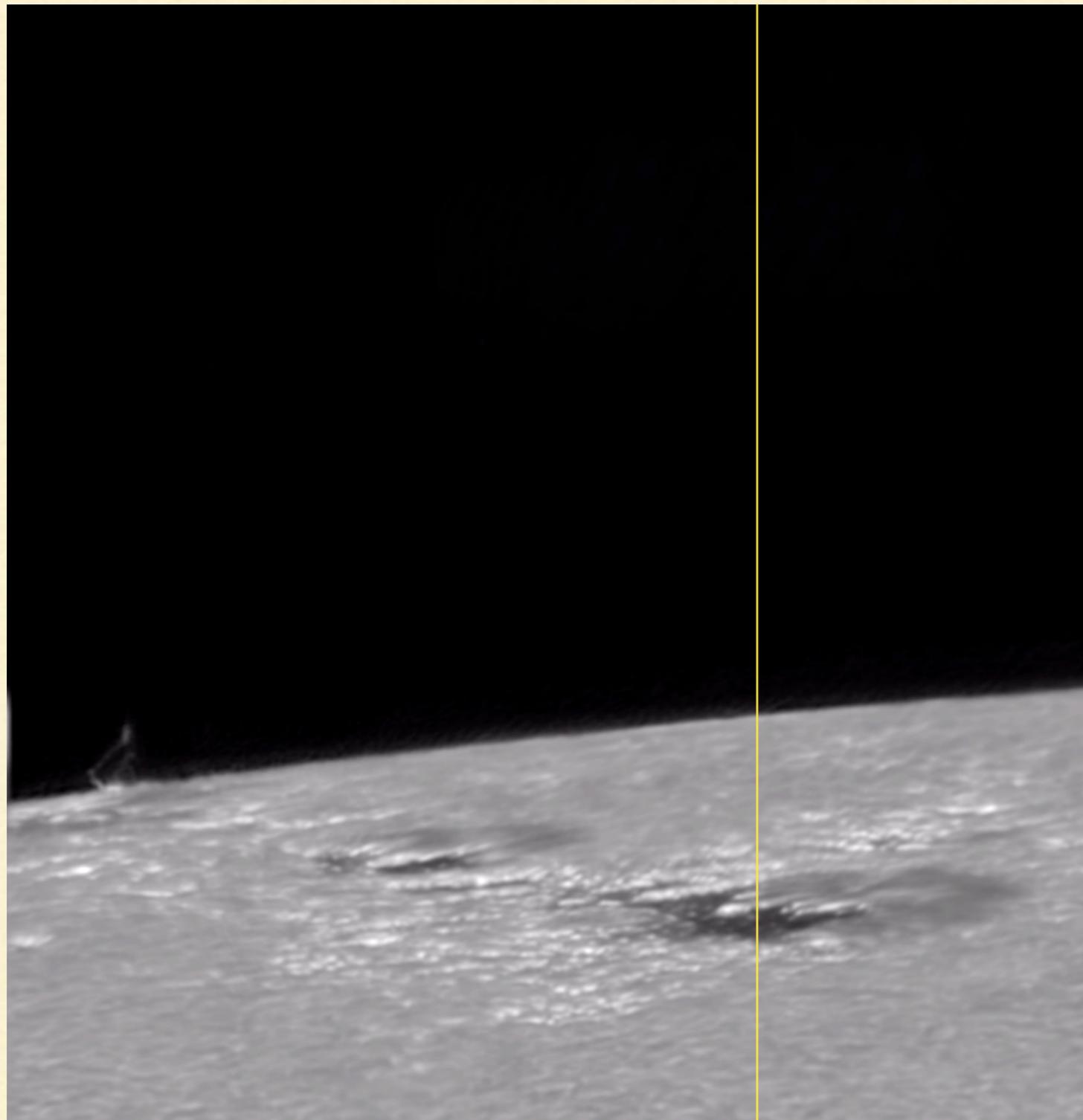


Observation Setup

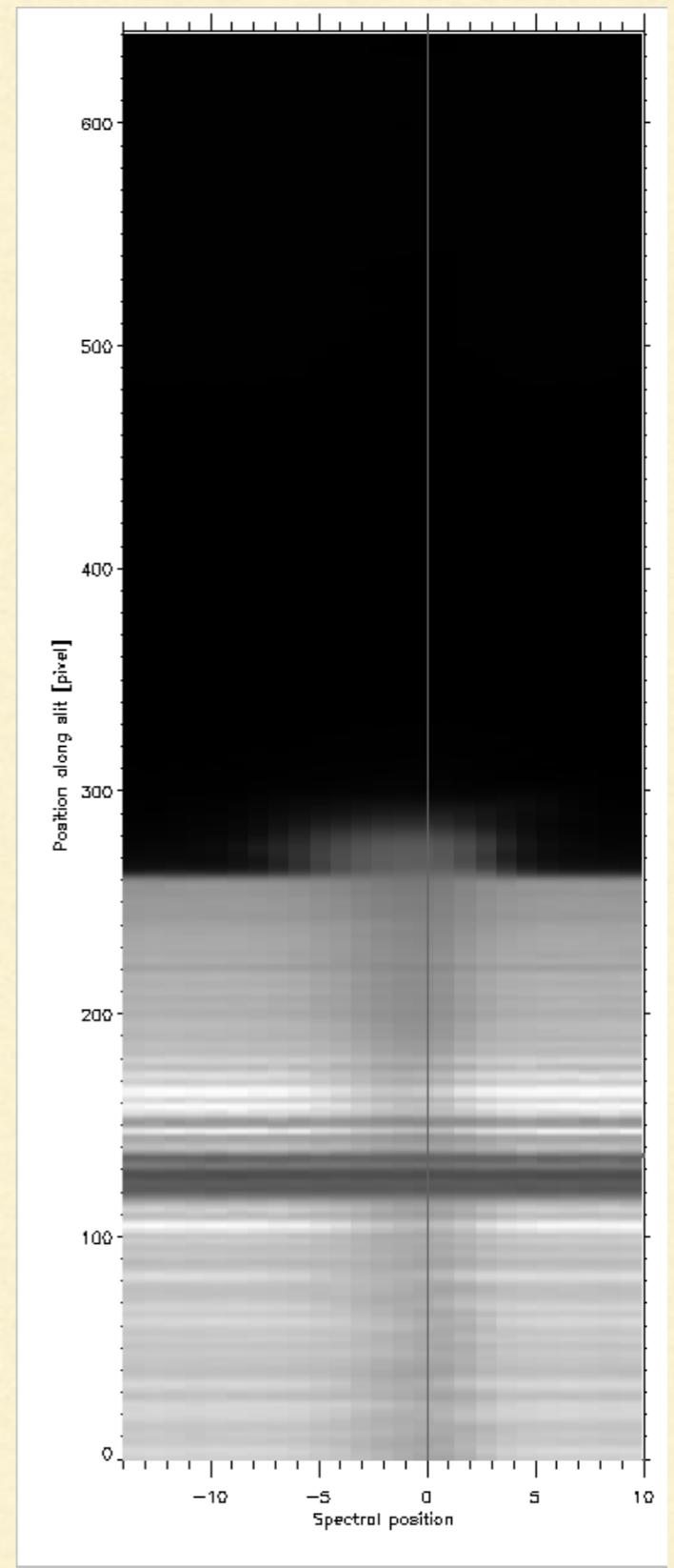


Credit: Mats Löfdahl

Limb Observation at 7772 Å



Using Crispex (Vissers et al. 2102)

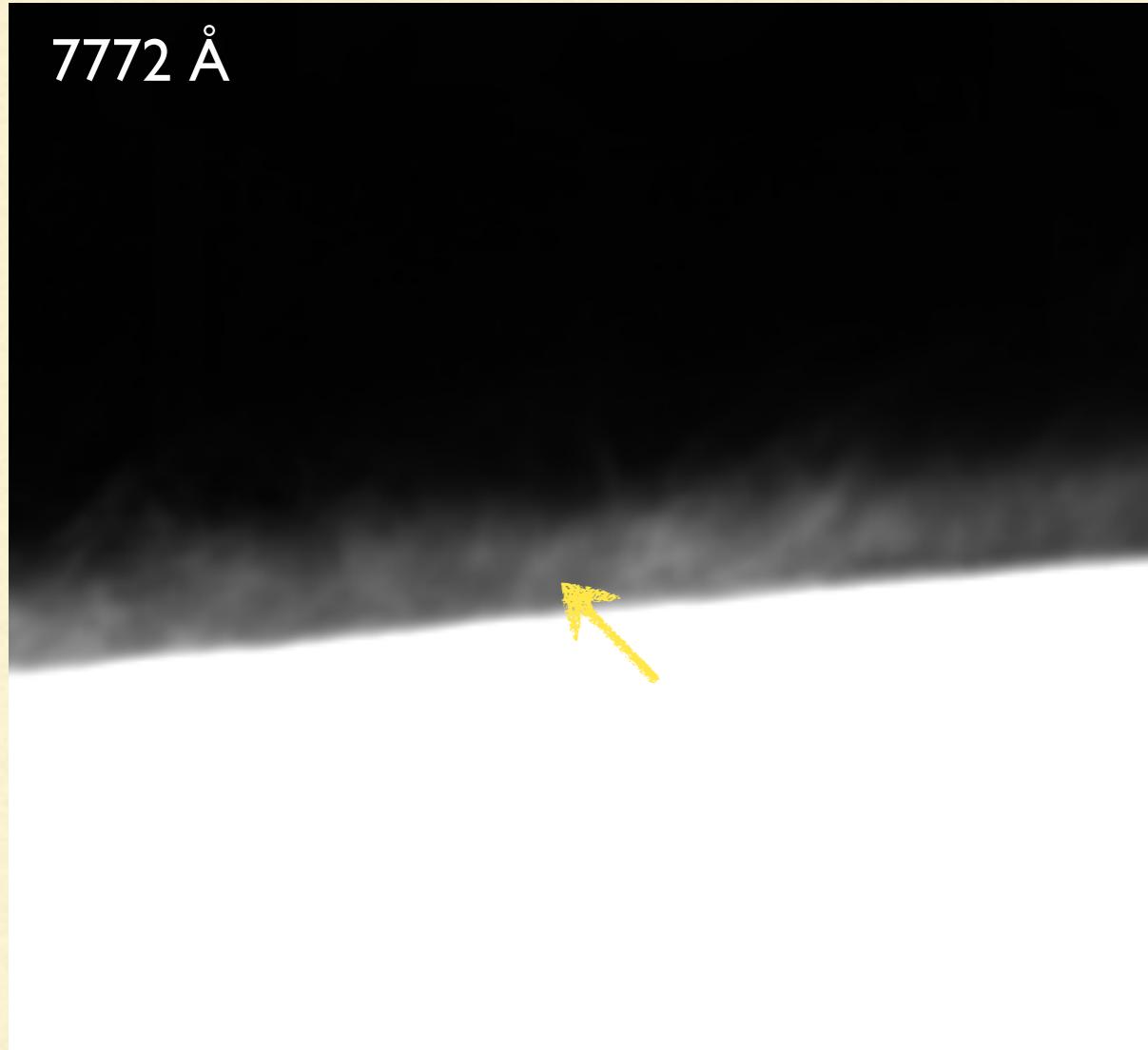


Observations (what is observable what is not!)

- Spicules?
- Dark Gap
- Prominence
- On disk observation
- Surge
- Filament

Spicules?

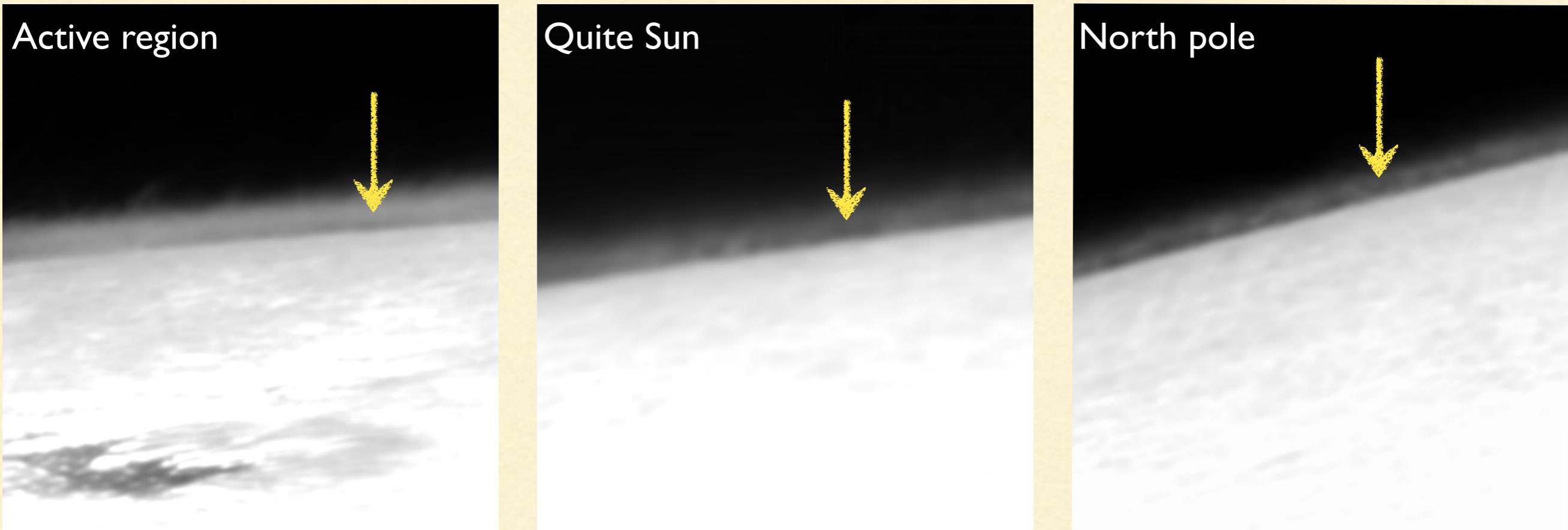
7772 Å



8542 Å

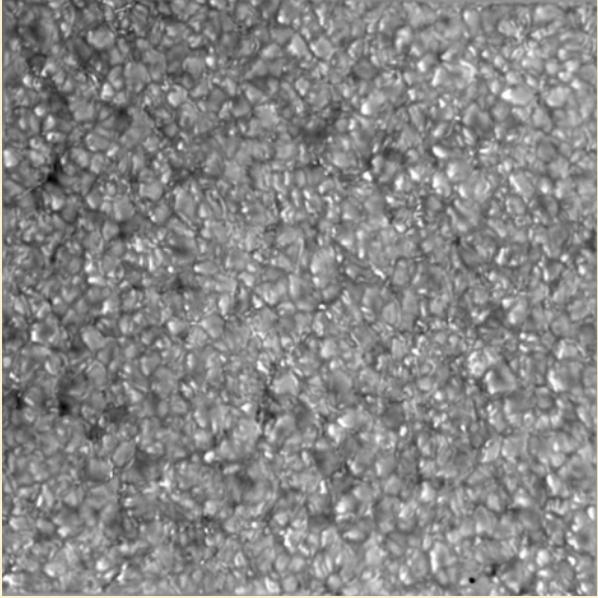


Dark Gap



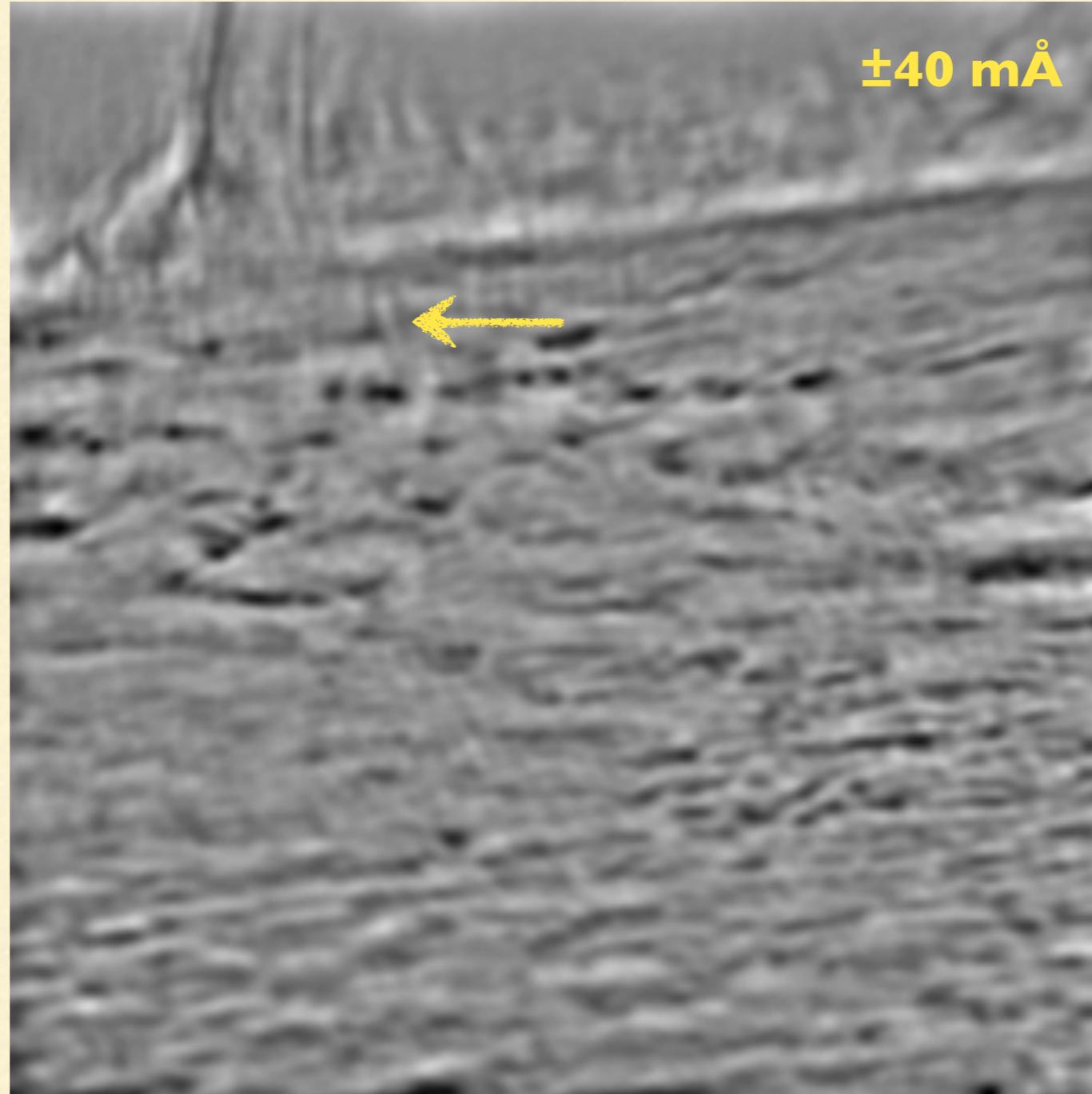
Prominence



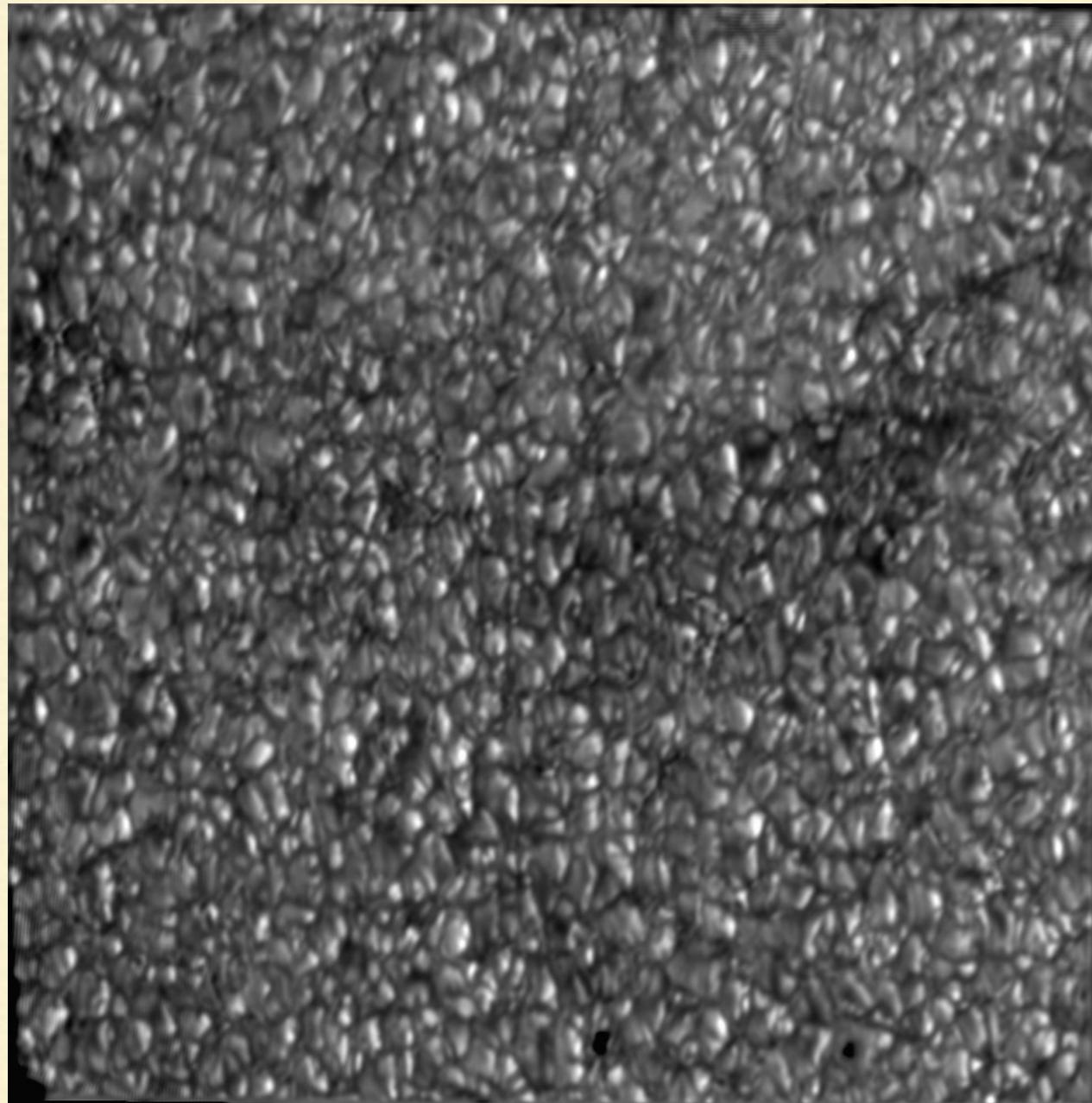


Close to disk center

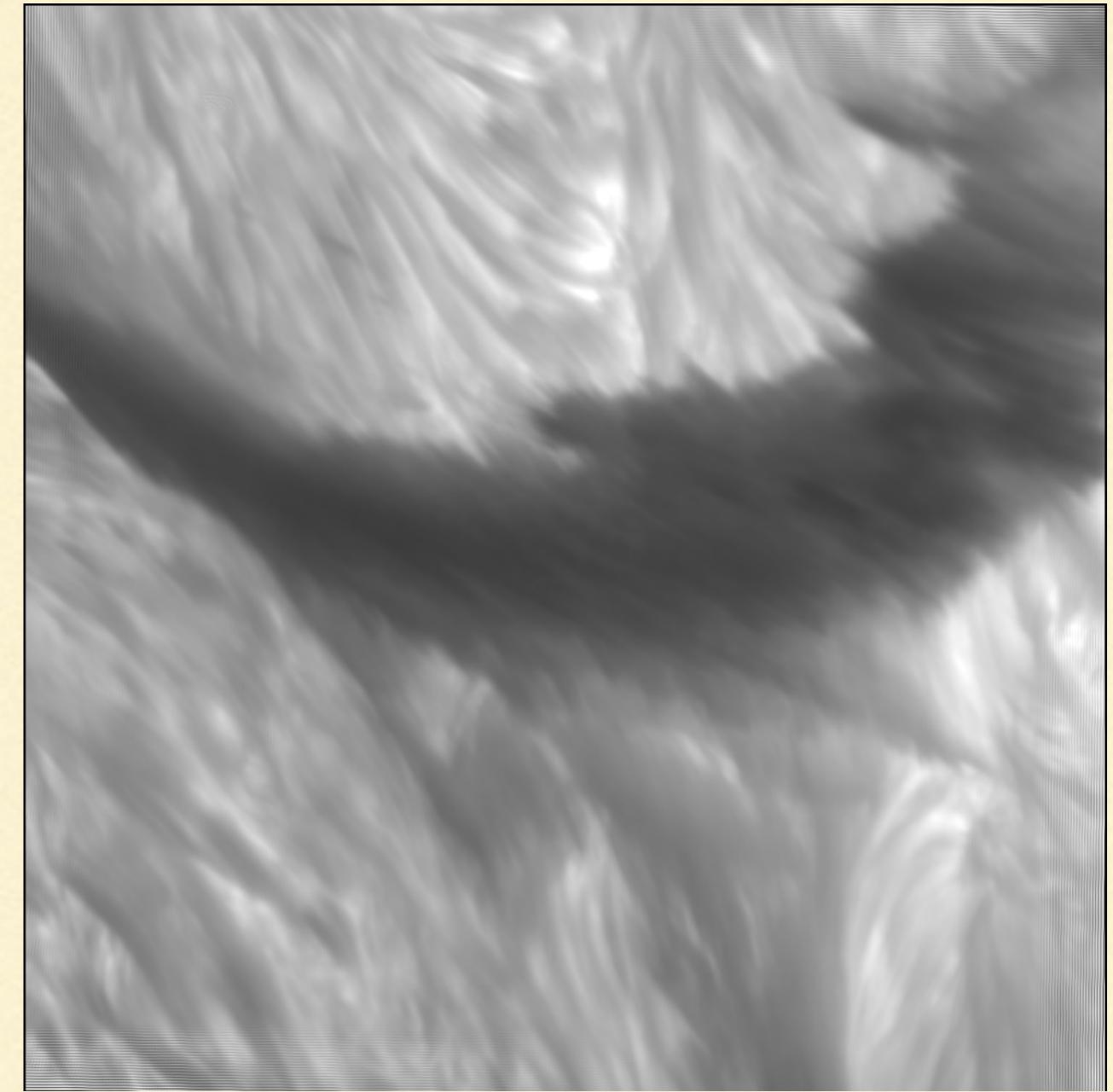
On Disk



Filament :(



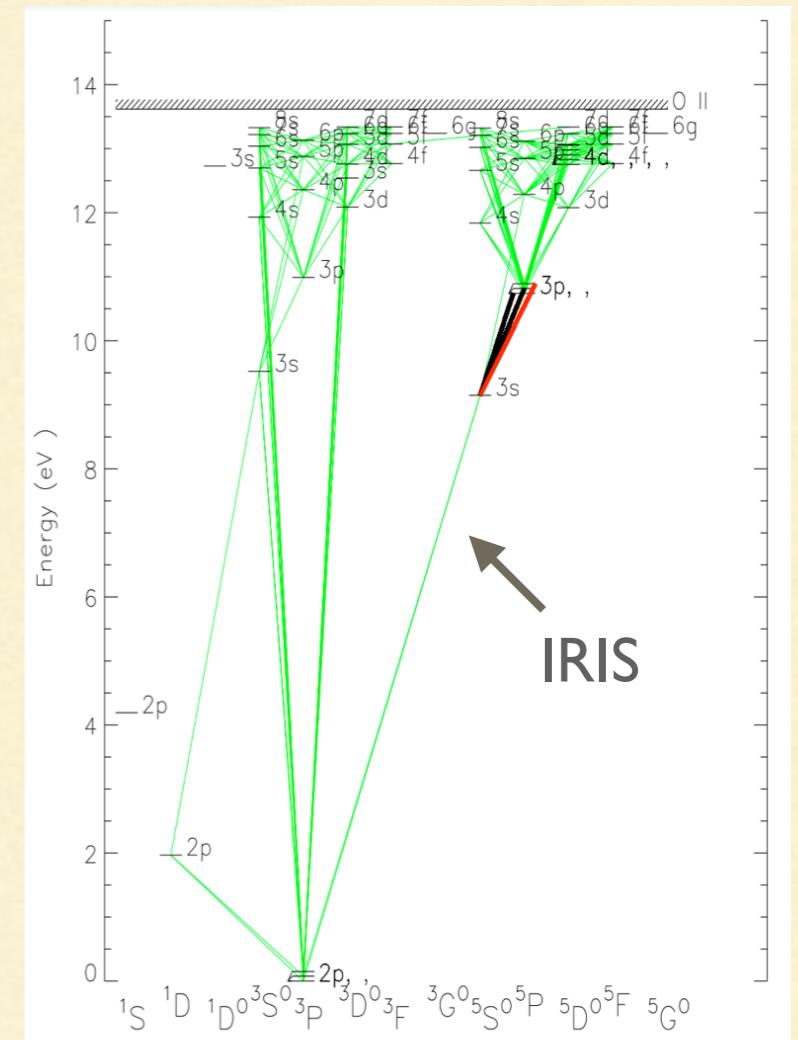
7772 Å

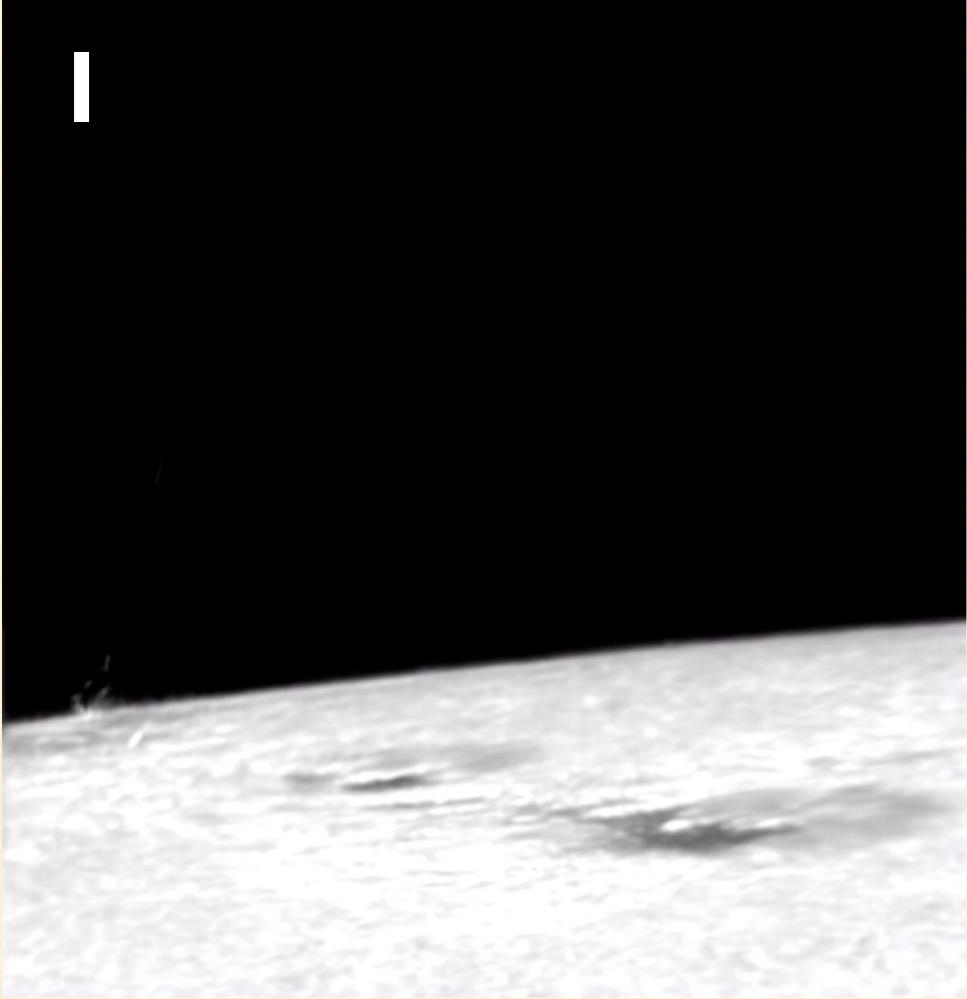


6563 Å (H α)

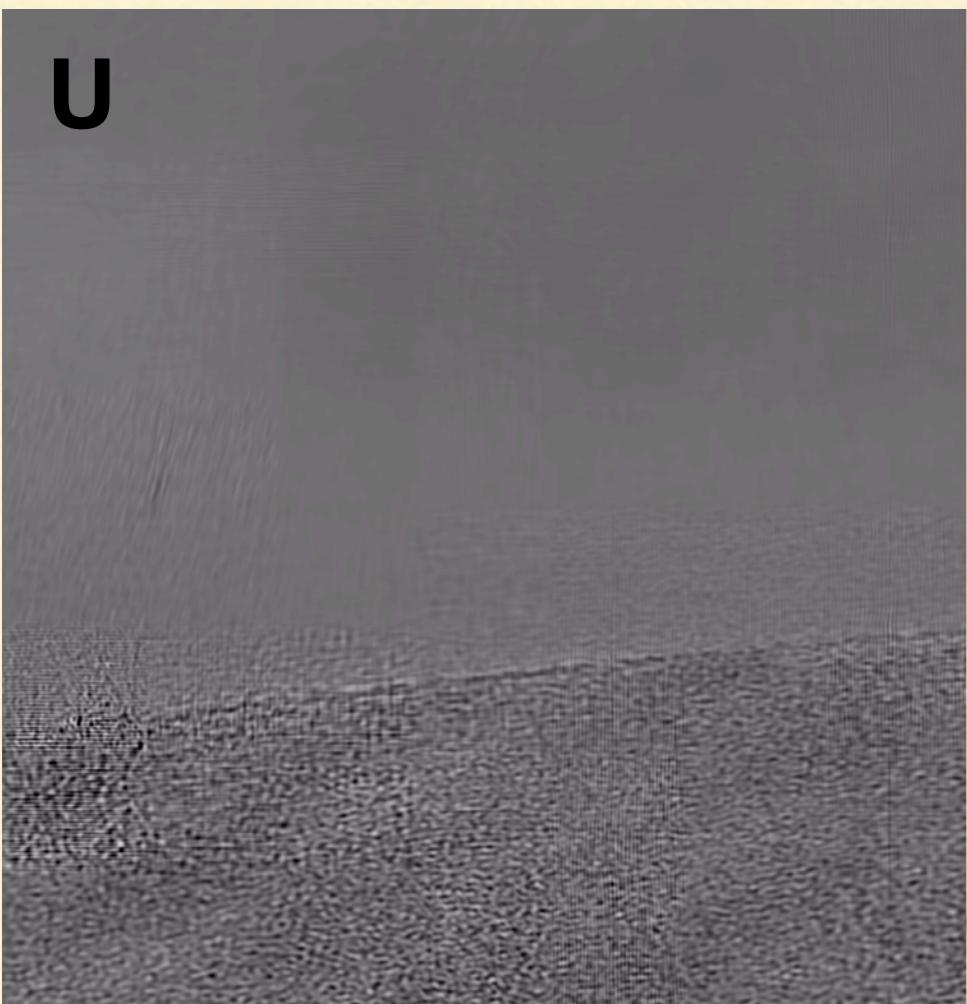
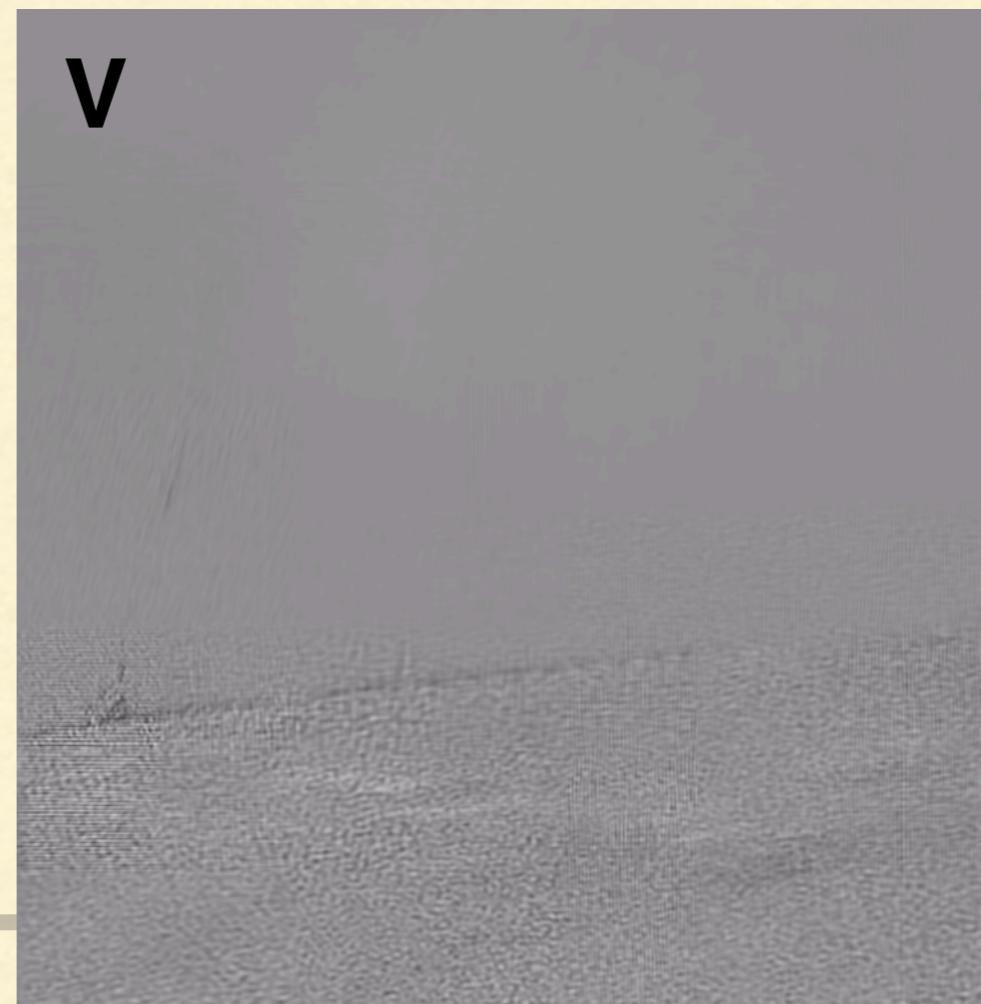
Future work

- Comparison with synthetic line profile
- Co-observation with IRIS
- Polarimetry data at Ca II (8542 Å)



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854.2 nm

Q**U****V**

Thanks!